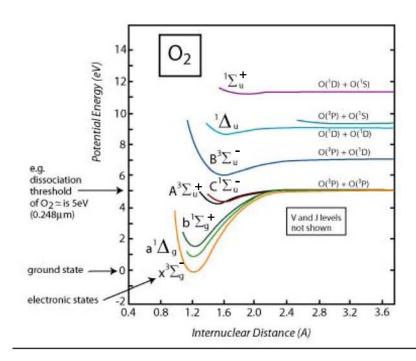
Energy Levels in Molecules, contd.

e.g. energy level diagram for O2



 $\begin{array}{c} \underline{\text{Conversions}} \colon \ 1\text{ev} = 8067 \ \text{cm}^{\text{-1}} = 1.24 \ \mu\text{m} \ (\text{near infrared}) \\ 2\text{ev} = 16134 \ \text{cm}^{\text{-1}} = 0.62 \ \mu\text{m} \\ 3\text{ev} = 24201 \ \text{cm}^{\text{-1}} = 0.41 \ \mu\text{m} \\ 5\text{ev} = 40335 \ \text{cm}^{\text{-1}} = 0.248 \ \mu\text{m} \ (\text{ultraviolet}) \\ \end{array}$

Absorption and emission by gases

For transitions between two states, initial i and final f, the Einstein "b" coefficient is given by

$$b_{fi} = \frac{1}{4\pi} \times \frac{8\pi^3 v_{fi}}{3hc} \left| \int \Phi_f^* \left(\hat{\mu} \Phi_i \right) dV \right|^2 \left| \int \sum_f^* \sum_i d\sigma \right|^2$$

and we need this to be \gg 0 for strong ("allowed") absorption or emission. This defines "selection rules".

 $\Phi(r, \theta, \phi)$ = space-dependent part of the wavefunction describing state of molecule.

$$\sum \left(\sigma\right) = \text{spin-dependent part of wavefunction}$$
 (total wavefunction = Φ , \sum)

 $\vec{\mu}$ = dipole moment operator = \vec{er}

(a) Selection rules from $\bar{\mu}$ (dipole)

For strong absorption/emission need $\hat{\mu} \neq 0$

- (i) <u>pure rotational transitions</u> permanent dipole <u>required</u>
- (ii) <u>vibrational transitions</u> where vibration does <u>not</u> produce a dipole moment permanent dipole <u>required</u>
- (iii) <u>vibrational transitions</u> where vibration produces a (transient) dipole moment permanent dipole <u>not required</u>.
- (iv) when vibrational transitions allowed then simultaneous vibrational + rotational transitions allowed
- (v) <u>electronic transitions</u> permanent dipole <u>not required</u> (always a dipole between nuclei and electrons)

(b) Selection rules from $\Phi_{\rm i}$ and $\Phi_{\rm f}$ (symmetry)

Since $\vec{\mu}$ (or \vec{r}) is an <u>odd</u> function in space, and we require total integrated $\Phi_f^* \vec{\mu} \Phi_i$ to be an <u>even</u> function to prevent significant cancellation between volume elements, then Φ_f^* and Φ_i must have opposite parity.

(La Porte's Rule). Formally this leads to:

i) for rotational energy transitions in a linear molecule

$$\Delta J = \pm 1$$

(<u>special case</u> is $\Delta J = 0, \pm 1$ for a non-diatomic simultaneous rotational-vibrational transition where vibration is \bot to axis)

ii) for vibrational energy transitions in all molecules

$$\Delta V = \pm 1$$
 (i.e. generally $V = 0 \rightarrow V = 1$)

(<u>i.e.</u> overtones and combinations are not allowed) (nv) $(v_1 + v_2)$

- (c) selection rules from \sum (spin)

For $b_{f_i} \neq 0$ need $\int \sum_i^* \sum_j d\sigma \neq 0$ but since $\sum_i^* s$ are orthogonal functions then need $\sum_i^* s = \sum_j^* s$

Hence, transitions cannot involve a change of spin

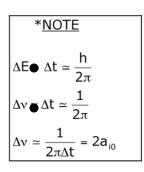
e.g. O₂ (see Energy Level diagram)

(* observed as very weak "magnetic dipole" transitions)

Shape of absorption and emission lines

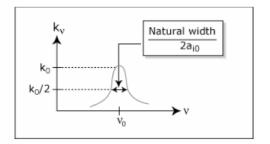
(a) "natural" broadening due to (Heisenberg*) uncertainty in energy of upper and lower levels.

Integrated line strength = $S_{i0} = \int_{0}^{\infty} k_{\nu} d\nu$



$$= \int_0^\infty \frac{k_0}{1 + \left(\frac{v - v_0}{a_{i0}}\right)^2} \ dv$$

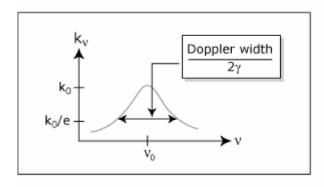




(recall
$$a_{i0} = \frac{1}{4\pi\tau_{i0}}$$
)
$$\uparrow \qquad (\Delta t)$$

(b) "doppler" broadening due to motion toward and away from observer (use Maxwellian distribution of velocities)

$$S_{i0} = \int_{0}^{\infty} k_{v} dv = \int_{0}^{\infty} \frac{S_{i0}}{\gamma \sqrt{\pi}} exp \left[-\left(\frac{v - v_{0}}{\gamma}\right)^{2} \right] dv$$



$$\left(\gamma = \frac{v_0}{c} \left(\frac{2KT}{m}\right)^{\frac{1}{2}}\right)$$

$$\left(\text{c.f. } v_{\text{molecule}} = \left(\frac{8 \text{ K T}}{\pi \text{m}}\right)^{\frac{1}{2}}\right)$$

(c) "pressure" broadening due to collisional and long distance interactions perturbing the energy levels

$$S_{i0} = \int k_{\nu} \, d\nu = \int\limits_{0}^{\infty} \left[\frac{k_{0}}{1 + \left(\frac{\nu - \nu_{0}}{\Gamma/4\pi} \right)^{2}} \right] \, d\nu$$

 $\Gamma \propto \text{collisional frequency} \blacktriangleleft$

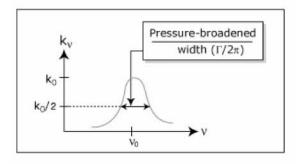
NOTE Collision freq.(sec⁻¹) in a 2-component gas $(m_1, r_1, n_1 \text{ and } m_2, r_2, n_2)$ is:

$$\left[(n_1 + n_2) (r_1 + r_2)^2 \left(8 \pi K T \left(\frac{1}{m_1} + \frac{1}{m_2} \right) \right)^{\frac{1}{2}} \right]$$

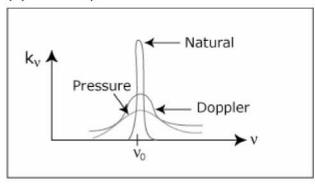
i.e.
$$\Gamma = \Gamma_0$$
 (1 atm, 273K) $P^{\alpha} T^{-\overline{\beta}}$ $\left(\overline{\alpha} \simeq 1, \ \overline{\beta} \simeq \frac{1}{2}\right) \ \left(P = n \, K \, T\right)$

 $\left(\Gamma_{0}, \stackrel{\leftarrow}{\alpha}, \stackrel{\rightarrow}{\beta} \text{ from laboratory measurements}\right)$

$$(k_0 = 4S_{i0}/\Gamma)$$



(d) Summary



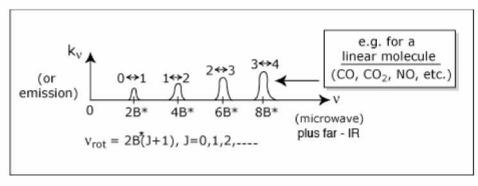
If increaseT: Natural → Doppler

If increaseP: Doppler → Pressure

Absorption bands

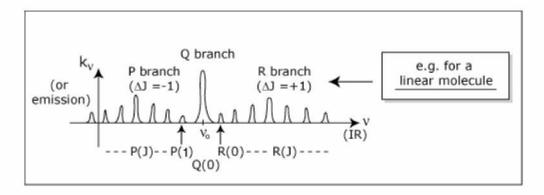
Can get rotational, vibrational, and electronic transitions together:

(a) rotational bands (e.g. H_2O) – due to $\Delta J = \pm 1$



12.815, Atmospheric Radiation Prof. Ronald Prinn

(b) vibrational-rotation bands – due to $\Delta J = 0, \pm 1$ and $\Delta v = \pm 1$



rotn. rise + vibn. rise
$$v_{vib-rot} = v_0 \begin{cases} +2B*(J+1) & J=0, 1, 2.....(R) \\ +0 & J=0, 1, 2....(Q) \\ -2B*(J) & J=1, 2, 3....(P) \end{cases}$$
 rotn. fall + vibn. rise

*B =
$$\frac{h^2}{8\pi^2 I}$$

(c) <u>electronic – vibration – rotation bands</u>

