

Lecture 8: Perfect Diamagnetism

Outline

1. **Description of a Perfect Diamagnet**
 - **Method I and Method II**
 - **Examples**
2. **Energy and Coenergy in Methods I and II**
3. **Levitating magnets and Maglev trains**

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Description of Perfect Diamagnetism

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Please see: Figure 2.17, page 50, from Orlando, T., and K. Delin.
Foundations of Applied Superconductivity. Reading, MA:
Addison-Wesley, 1991. ISBN: 0201183234.

Surface currents or internal induced magnetization?

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Methods I and II

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$$\mathbf{B} = \mu_0 \mathbf{H}^I \qquad \mathbf{B} = \mu_0 (\mathbf{H}^{II} + \mathbf{M})$$

Method I

Method II

B field is the same in both methods.

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Example: Magnetized Sphere

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$$\mathbf{M} = M_0 \mathbf{i}_z = M_0 (\cos \theta \mathbf{i}_r - \sin \theta \mathbf{i}_\theta) \qquad \text{for } r \leq R.$$

$$\left. \begin{array}{l} \nabla \times \mathbf{H} = \mathbf{J} = 0 \\ \nabla \cdot \mathbf{B} = 0 \\ \mathbf{B} = \mu_0 (\mathbf{H} + \mathbf{M}) \end{array} \right\} \begin{array}{l} \longrightarrow \mathbf{H} = -\nabla \psi \\ \longrightarrow \nabla^2 \psi = \nabla \cdot \mathbf{M} \end{array}$$

For this example: $\nabla^2 \psi = 0$ Laplace's equation

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Boundary Conditions

Inside the sphere:

$$\psi(r \leq R) = C_1 r \cos \theta \quad \Rightarrow \quad \mathbf{H}(r \leq R) = -C_1 (\cos \theta \mathbf{i}_r - \sin \theta \mathbf{i}_\theta)$$

Outside the sphere:

$$\psi(r \geq R) = C_2 \frac{\cos \theta}{r^2} \quad \Rightarrow \quad \mathbf{H}(r \geq R) = \frac{C_2}{r^3} (2 \cos \theta \mathbf{i}_r + \sin \theta \mathbf{i}_\theta)$$

Boundary Conditions:

$$\mathbf{n} \times (\mathbf{H}_2 - \mathbf{H}_1) = \mathbf{K} \quad \Rightarrow \quad C_1 R^3 = C_2$$

$$\mathbf{n} \cdot (\mathbf{B}_2 - \mathbf{B}_1) = 0 \quad \Rightarrow \quad \mathbf{i}_r \cdot (\mathbf{H}|_{r=R^+} - \mathbf{H}|_{r=R^-}) = \mathbf{i}_r \cdot \mathbf{M}|_{r=R^-}$$

$$\text{Therefore, } \mathbf{M} = M_o (\cos \theta \mathbf{i}_r - \sin \theta \mathbf{i}_\theta) \quad \Rightarrow \quad C_1 = \frac{M_o}{3}$$

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Magnetized Sphere

$$\mathbf{H}(r \leq R) = -\frac{M_o}{3} \mathbf{i}_z = -\frac{M_o}{3} (\cos \theta \mathbf{i}_r - \sin \theta \mathbf{i}_\theta)$$

$$\mathbf{H}(r \geq R) = \frac{M_o}{3} \left(\frac{R}{r}\right)^3 (2 \cos \theta \mathbf{i}_r + \sin \theta \mathbf{i}_\theta)$$

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Magnetized Sphere in field

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$$\mathbf{H}(r \leq R) = \left(H_o - \frac{M_o}{3} \right) \mathbf{i}_z$$

$$\mathbf{H}_{\text{app}} = H_o \mathbf{i}_z \quad \mathbf{H}(r \geq R) = \left(H_o + \frac{2}{3} M_o \left(\frac{R}{r} \right)^3 \right) \cos \theta \mathbf{i}_r - \left(H_o - \frac{1}{3} M_o \left(\frac{R}{r} \right)^3 \right) \sin \theta \mathbf{i}_\theta$$

For this to describe a superconductor (bulk limit), then $\mathbf{B}=0$ inside.

$$\text{Therefore, } 0 = \mathbf{M} + \mathbf{H}(r \leq R) = M_o \mathbf{i}_z + \left(H_o - \frac{M_o}{3} \right) \mathbf{i}_z$$

$$\text{So that } M = -\frac{3}{2} H_o$$

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Comparison of Methods

Method I

$$\mathbf{B}^I = \mu_o \mathbf{H}^I$$

Inside a bulk superconducting sphere:

$$\mathbf{B}^I = 0$$

$$\mathbf{H}^I = 0$$

$$\mathbf{K}^I = -\text{Re} \left\{ \frac{3}{2} \hat{H}_o \sin \theta \right\} \mathbf{i}_\phi$$

Method II

$$\mathbf{B} = \mu_o (\mathbf{H}^{II} + \mathbf{M})$$

$$\mathbf{B}^{II} = 0$$

$$\mathbf{H}^{II} = \frac{3}{2} H_o \mathbf{i}_z$$

$$\mathbf{M} = -\frac{3}{2} H_o \mathbf{i}_z$$

$$\mathbf{K}^{II} = 0$$

B field is the same, but not **H**.

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Methods I and II: Summary

Maxwell's Equations

Method I	$\nabla \cdot \mathbf{B} = 0$	Method II
$\nabla \times \mathbf{H}^I = \mathbf{J}^I$		$\nabla \times \mathbf{H}^{II} = \mathbf{J}^{II}$
$\mathbf{B} = \mu_0 \mathbf{H}^I$		$\mathbf{B} = \mu_0 (\mathbf{H}^{II} + \mathbf{M})$
$\nabla \times \mathbf{B} = \mu_0 \mathbf{J}^I$		$\nabla \times \mathbf{B} = \mu_0 \mathbf{J}^{II} + \mu_0 \underbrace{\nabla \times \mathbf{M}}_{\mathbf{J}_{s,ind}}$
$\mathbf{J}^I = \underbrace{(\mathbf{J}_{s,app} + \mathbf{J}_{s,ind})}_{\mathbf{J}_s^I} + \mathbf{J}_n$		$\mathbf{J}^{II} = \mathbf{J}_{s,app} + \mathbf{J}_n$

London Equations

$\mathbf{E} = \frac{\partial}{\partial t} (\wedge \mathbf{J}_s^I)$	$\mathbf{E} = \frac{\partial}{\partial t} (\wedge \mathbf{J}_s^{II}) + \frac{\partial}{\partial t} (\wedge (\nabla \times \mathbf{M}))$
$\nabla \times (\wedge \mathbf{J}_s^I) = -\mathbf{B}$	$\nabla \times (\wedge \mathbf{J}_s^{II}) + \nabla \times (\wedge (\nabla \times \mathbf{M})) = -\mathbf{B}$

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Why Method II

Question:

The constitutive relations and London's Equations have gotten much more difficult. So why do Method II?

Answer:

The Energy and Thermodynamics are easier, especially when there is no applied current.

So we will find the energy stored in both methods.

Poynting's theorem is a result of Maxwell's equation, so both methods give

$$-\oint_{\Sigma} (\mathbf{E} \times \mathbf{H}) \cdot d\mathbf{s} = \int_V \left(\mathbf{E} \cdot \frac{\partial \mathbf{D}}{\partial t} + \mathbf{H} \cdot \frac{\partial \mathbf{B}}{\partial t} + \mathbf{E} \cdot \mathbf{J} \right) dv$$

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Method I: The Energy

Combining the constitutive relations with Poynting's theorem,

$$-\oint_{\Sigma} (\mathbf{E} \times \mathbf{H}^I) \cdot d\mathbf{s} = \int_V \underbrace{\left(\mathbf{E} \cdot \frac{\partial \mathbf{D}}{\partial t} + \mathbf{H}^I \cdot \frac{\partial \mathbf{B}}{\partial t} + \mathbf{J}_s^I \cdot \frac{\partial (\wedge \mathbf{J}_s^I) }{\partial t} \right)}_{\text{Power } dW/dt \text{ in the E\&M field}} dv + \int_V \mathbf{E} \cdot \mathbf{J}_n dv$$

The energy stored in the electromagnetic field is

$$dW = \int_V \left(\mathbf{E} \cdot d\mathbf{D} + \mathbf{H}^I \cdot d\mathbf{B} + \mathbf{J}_s^I \cdot d(\wedge \mathbf{J}_s^I) \right) dv$$

So that the energy W is a function of \mathbf{D} , \mathbf{B} , and $\wedge \mathbf{J}_s^I = \mathbf{v}_s \cdot$.
However, one rarely has control over these variables, but rather over their conjugates \mathbf{E} , \mathbf{H}^I , and \mathbf{J}_s^I .



Method I: The Coenergy

The coenergy is a function of \mathbf{E} , \mathbf{H}^I , and \mathbf{J}_s^I is defined by

$$W + \widetilde{W} = \int_V \left(\mathbf{E} \cdot \mathbf{D} + \mathbf{H}^I \cdot \mathbf{B} + \mathbf{J}_s^I \cdot (\wedge \mathbf{J}_s^I) \right) dv$$

and with

$$dW = \int_V \left(\mathbf{E} \cdot d\mathbf{D} + \mathbf{H}^I \cdot d\mathbf{B} + \mathbf{J}_s^I \cdot d(\wedge \mathbf{J}_s^I) \right) dv$$

gives

$$\leftarrow \text{EQS} \rightarrow \quad \leftarrow \text{MQS} \rightarrow$$

$$d\widetilde{W} = \int_V \left(\mathbf{D} \cdot d\mathbf{E} + \mathbf{B} \cdot d\mathbf{H}^I + \wedge \mathbf{J}_s^I \cdot d\mathbf{J}_s^I \right) dv$$

(The coenergy is the Free Energy at zero temperature.)



Interpretation of the Coenergy

Consider the case where there are only magnetic fields:

$$\tilde{W} = \int_{\mathbf{H}^I} d\tilde{W} = \int_V \mathbf{B} \cdot d\mathbf{H}^I dv \quad W = \int_{\mathbf{B}} dW = \int_{\mathbf{B}} \int_V \mathbf{H}^I \cdot d\mathbf{B} dv$$

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Please see: Figure 4.12, page 184, from Orlando, T., and K. Delin.

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$$W_m + \tilde{W}_m = H_f B_f$$

The energy and coenergy contain the same information.

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Method II

$$dW = \int_V (\mathbf{E} \cdot d\mathbf{D} + \mathbf{H}^{II} \cdot d\mathbf{B} + \mathbf{J}_S^{II} \cdot d(\wedge \mathbf{J}_S^{II} + \wedge \nabla \times \mathbf{M})) dv$$

$$d\tilde{W} = \int_V (\mathbf{D} \cdot d\mathbf{E} + \mathbf{B} \cdot d\mathbf{H}^{II} + (\wedge \mathbf{J}_S^{II} + \wedge \nabla \times \mathbf{M}) \cdot d\mathbf{J}_S^{II}) dv$$

In the important case when of the MQS limit and $\mathbf{J}_S^{II} = \mathbf{J}_{s,app} = 0$

$$dW|_{\mathbf{J}_{s,app}=0} = \int_V (\mathbf{H}^{II} \cdot d\mathbf{B}) dv$$

$$d\tilde{W}|_{\mathbf{J}_{s,app}=0} = \int_V (\mathbf{B} \cdot d\mathbf{H}^{II}) dv$$

Note that these two relations apply also in free space.

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Example: Energy of a Superconducting Sphere

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 Addison-Wesley, 1991. ISBN: 0201183234.

$$\mathbf{B}(r \leq R) = 0 \quad \mathbf{H}^{II}(r < R) = \frac{3}{2} H_0 \mathbf{i}_z$$

$$\mathbf{H}(r \geq R) = \operatorname{Re} \left\{ \hat{H}_0 \left(1 - \left(\frac{R}{r} \right)^3 \right) \cos \theta \right\} \mathbf{i}_r - \operatorname{Re} \left\{ \hat{H}_0 \left(1 + \frac{1}{2} \left(\frac{R}{r} \right)^3 \right) \sin \theta \right\} \mathbf{i}_\theta$$

$$W_{inside} = \int_B \int_V (\mathbf{H}^{II} \cdot d\mathbf{B}) dv = 0$$

$$W_{outside} = \int_B \int_V (\mathbf{H}^{II} \cdot d\mathbf{B}) dv = \int_V \frac{1}{\mu_0} \mathbf{B}^2 dv \neq 0$$

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Magnetic Levitation

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Magnetic vs. Gravitational Forces

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Magnetic Levitation Equilibrium Point

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$$W_m = \int_V \frac{1}{\mu_0} B^2 dv \approx \frac{\pi R^2 \eta}{\mu_0} B^2 = \frac{\Phi^2}{8\pi\mu_0\eta}$$

$$\Phi = \pi R^2 B_0 = 2\pi R\eta B$$

$$f_m = -\frac{\partial}{\partial \eta} W_m(\Phi, \eta) \approx \frac{\Phi^2}{8\pi\mu_0\eta^2} \quad \text{Force is upwards}$$

Equilibrium point $f_m(\eta_0) = mg$  $\eta_0 = \sqrt{\frac{\Phi^2}{8\pi\mu_0 mg}}$

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Levitating magnets and trains

$$f_m \approx \frac{\Phi^2}{8\pi\mu_0\eta^2}$$

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B = 1 Tesla
Area = 0.5 cm²
 $\eta_0 = 1$ cm
Force = 1 Newton

Enough to lift magnet,
but not a train

B = 2 Tesla
Area = 100 cm²
 $\eta_0 = 10$ cm
Force = 1,200 Newtons

Enough to lift a train

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Maglev Train

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Magnets on the train are superconducting magnets;
the rails are ohmic!

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Principle of Maglev

The train travels at a velocity U , and the moving flux lines and the rails “see” a moving magnetic field at a frequency of $\omega \sim U/R$.

If this frequency is much larger than the inverse of the magnetic diffusion time,

$$\tau_m = \mu\sigma_o R d$$

then the flux lines are “repelled” from the ohmic rails.

$$\omega\tau_m \gg 1 \qquad U \gg \frac{1}{\mu\sigma_o d}$$

From the previous numbers, $U > 40$ km/hr for levitation.

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Real trains have wheels and high-voltage rails

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Synchronous motor action down the rails provides thrust to accelerate train to the needed velocity to levitate, and provides a source of energy to further accelerate the train and to overcome the losses due to drag from the wind.

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Our Approach to Superconductivity

