

MATH 18.152 - OPTIONAL BONUS PROBLEM

18.152 Introduction to PDEs, Fall 2011

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Optional Bonus Problem, **Due: at the start of class on 11-29-11**

I. Consider the *Morawetz vectorfield* \bar{K}^μ on R^{1+3} defined by

$$(0.0.1) \quad \bar{K}^0 = 1 + t^2 + (x^1)^2 + (x^2)^2 + (x^3)^2,$$

$$(0.0.2) \quad \bar{K}^j = 2tx^j, \quad (j = 1, 2, 3).$$

a) Show that \bar{K} is future-directed and timelike. Above, (t, x^1, x^2, x^3) are the standard coordinates on R^{1+3} .

b) Show that

$$(0.0.3) \quad \partial_\mu \bar{K}_\nu + \partial_\nu \bar{K}_\mu = 4tm_{\mu\nu}, \quad (\mu, \nu = 0, 1, 2, 3),$$

where $m_{\mu\nu}$ denotes the Minkowski metric.

Remark 0.0.1. \bar{K} is said to be a *conformal Killing field* of the Minkowski metric because the right-hand side of (0.0.3) is proportional to $m_{\mu\nu}$.

c) Show that

$$(0.0.4) \quad m_{\mu\nu} T^{\mu\nu} = -(m^{-1})^{\alpha\beta} \partial_\alpha \phi \partial_\beta \phi,$$

where $T_{\mu\nu} \stackrel{\text{def}}{=} \partial_\mu \phi \partial_\nu \phi - \frac{1}{2} m_{\mu\nu} (m^{-1})^{\alpha\beta} \partial_\alpha \phi \partial_\beta \phi$ is the energy-momentum tensor corresponding to the linear wave equation, and $T^{\mu\nu} \stackrel{\text{def}}{=} (m^{-1})^{\mu\alpha} (m^{-1})^{\nu\beta} T_{\alpha\beta}$ is the energy-momentum tensor with its indices raised.

d) Show that $\partial_\mu (\bar{K}) J^\mu = 2t(m^{-1})^{\alpha\beta} \partial_\alpha \phi \partial_\beta \phi$ whenever ϕ is a C^2 solution to the linear wave equation $(m^{-1})^{\mu\nu} \partial_\mu \partial_\nu \phi = 0$, where

$$(0.0.5) \quad (\bar{K}) J^\mu \stackrel{\text{def}}{=} -T^{\mu\nu} \bar{K}_\nu.$$

e) Show that $\partial_\mu \tilde{J}^\mu = 0$ whenever ϕ is a C^2 solution to the linear wave equation $(m^{-1})^{\mu\nu} \partial_\mu \partial_\nu \phi = 0$, where

$$(0.0.6) \quad \tilde{J}^\mu \stackrel{\text{def}}{=} (\bar{K}) J^\mu - 2t\phi(m^{-1})^{\mu\alpha} \partial_\alpha \phi + \phi^2(m^{-1})^{\mu\alpha} \partial_\alpha t.$$

f) Show that

$$(0.0.7) \quad (\bar{K}) J^0 = \frac{1}{4} \left\{ [1 + (t+r)^2] (\nabla_L \phi)^2 + [1 + (t-r)^2] (\nabla_{\underline{L}} \phi)^2 + 2[1 + t^2 + r^2] m^{\mu\nu} \partial_\mu \phi \partial_\nu \phi \right\}.$$

Above, $(m^{-1})^{\mu\nu} = -\frac{1}{2}L^\mu \underline{L}^\nu - \frac{1}{2}\underline{L}^\mu L^\nu + \not{n}^{\mu\nu}$ is the standard null decomposition of $(m^{-1})^{\mu\nu}$ from class. In particular, $L^\mu = (1, \frac{x^1}{r}, \frac{x^2}{r}, \frac{x^3}{r})$, $\underline{L}^\mu = (1, -\frac{x^1}{r}, -\frac{x^2}{r}, -\frac{x^3}{r})$, $\nabla_L \phi = \partial_t \phi + \partial_r \phi$, $\nabla_{\underline{L}} \phi = \partial_t \phi - \partial_r \phi$, and $\not{n}^{\mu\nu} \partial_\mu \phi \partial_\nu \phi$ is the square of the Euclidean norm of the *angular* derivatives of ϕ . Here, $r \stackrel{\text{def}}{=} \sqrt{(x^1)^2 + (x^2)^2 + (x^3)^2}$ denotes the standard spherical coordinate on \mathbb{R}^3 , and ∂_r denotes the standard radial derivative.

Hint: The following expansions in terms of L and \underline{L} may be very helpful:

(0.0.8)

$$\bar{K}^\mu = \frac{1}{2} \left\{ [1 + (r+t)^2] L^\mu + [1 + (r-t)^2] \underline{L}^\mu \right\},$$

(0.0.9)

$$(1, 0, 0, 0) = \frac{1}{2} (L^\mu + \underline{L}^\mu),$$

(0.0.10)

$$\begin{aligned} {}^{(\bar{K})} J^0 &= T(\bar{K}, \frac{1}{2}(L + \underline{L})) \\ &= \frac{1}{4} \left\{ [1 + (r+t)^2] T(L, L) + [1 + (r-t)^2] T(\underline{L}, \underline{L}) + ([1 + (r+t)^2] + [1 + (r-t)^2]) T(L, \underline{L}) \right\}. \end{aligned}$$

f)

Show that

(0.0.11)

$$\tilde{J}^0 = \frac{1}{4} \left\{ [1 + (t+r)^2] (\nabla_L \phi)^2 + [1 + (t-r)^2] (\nabla_{\underline{L}} \phi)^2 + 2[1 + t^2 + r^2] \not{n}^{\mu\nu} \partial_\mu \phi \partial_\nu \phi \right\} + 2t\phi \partial_t \phi - \phi^2.$$

g) Show that

(0.0.12)

$$-\int_{\mathbb{R}^3} \phi^2 d^3x = \frac{2}{3} \int_{\mathbb{R}^3} r\phi \partial_r \phi d^3x$$

whenever ϕ is a C^1 , compactly supported function.

Hint: Use the identity $1 = \partial_r r$ together with integration by parts in spherical coordinates and the fact that $d^3x = r^2 \sin \theta dr d\theta d\phi$ in spherical coordinates.

h) Use parts f) and g) to show that

(0.0.13)

$$\tilde{J}^0 = \frac{1}{4} \left\{ (\nabla_L \phi)^2 + \left(\nabla_L [(t+r)\phi] \right)^2 + (\nabla_{\underline{L}} \phi)^2 + \left(\nabla_{\underline{L}} [(t-r)\phi] \right)^2 + 2[1 + t^2 + r^2] \not{n}^{\mu\nu} \partial_\mu \phi \partial_\nu \phi \right\}.$$

i) Finally, with the help of the vectorfield \tilde{J}^μ , part e), and part h), apply the divergence theorem on an appropriately chosen spacetime region to derive the following conservation law for smooth solutions to the linear wave equation $(m^{-1})^{\mu\nu} \partial_\mu \partial_\nu \phi = 0$:

(0.0.14)

$$\begin{aligned} & \frac{1}{4} \int_{\mathbb{R}^3} \left\{ (\nabla_L \phi)^2 + \left(\nabla_L [(t+r)\phi] \right)^2 + (\nabla_{\underline{L}} \phi)^2 + \left(\nabla_{\underline{L}} [(t-r)\phi] \right)^2 + 2[1+t^2+r^2] \not{m}^{\mu\nu} \partial_\mu \phi \partial_\nu \phi \right\} d^3x \\ &= \frac{1}{4} \int_{\mathbb{R}^3} \left\{ (\nabla_L \phi)^2 + \left(\nabla_L [(t+r)\phi] \right)^2 + (\nabla_{\underline{L}} \phi)^2 + \left(\nabla_{\underline{L}} [(t-r)\phi] \right)^2 + 2[1+t^2+r^2] \not{m}^{\mu\nu} \partial_\mu \phi \partial_\nu \phi \right\} d^3x \Big|_{t=0}, \end{aligned}$$

where the left-hand side is evaluated at time t , and right-hand side is evaluated at time $t = 0$. For simplicity, at each fixed t , you may assume that there exists an $R > 0$ such that $\phi(t, x)$ vanishes whenever $|x| \geq R$.

Remark 0.0.2. Note that the right-hand side of (0.0.14) can be computed in terms of the initial data alone. Note also that the different *null derivatives* of ϕ appearing on the left-hand side of (0.0.14) carry different weights. In particular, $\nabla_L \phi$ and the angular derivatives of ϕ have larger weights than $\nabla_{\underline{L}} \phi$. These larger weights are strongly connected to the following fact, whose full proof requires additional methods going beyond this course: $\nabla_L \phi$ and the angular derivatives of ϕ decay faster in t compared to $\nabla_{\underline{L}} \phi$.

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