

MATH 18.152 COURSE NOTES - CLASS MEETING # 11

18.152 Introduction to PDEs, Fall 2011

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Class Meeting # 11: The Method of Spherical Means

1. 1 + 3 SPACETIME DIMENSIONS AND THE METHOD OF SPHERICAL MEANS

We would now like to derive an analog of d'Alembert's formula in the physically relevant case of 1 + 3 dimensions. As we will see, the analogous formula, known as Kirchhoff's formula, can be derived through the following steps.

- Given a solution $u(t, x)$ to the 1 + 3 dimensional wave equation, we will define a spherical average of u centered at x . The average will depend on the averaging radius r .
- For fixed x , we will show that a slight modification of the average will solve the 1 + 1 dimensional wave equation in the unknowns (t, r) . With the help of our corollary to d'Alembert's formula, we will be able to find an explicit formula for this modified function.
- We will take a limit as the averaging goes to 0 in order to recover an expression for $u(t, x)$.

This procedure is known as the *method of spherical means*. The final result will be stated and proved as a theorem. Before proving the theorem, we will develop some preliminary estimates. We will use spherical coordinates $(r, \theta, \phi) \in [0, \infty) \times [0, \pi) \times [0, 2\pi)$ on \mathbb{R}^3 . Recall that if the spherical coordinates are centered at the Cartesian point (p^1, p^2, p^3) , then the standard Cartesian coordinates (x^1, x^2, x^3) are connected to spherical coordinates by

$$(1.0.1a) \quad x^1 = p^1 + r \sin \theta \cos \phi,$$

$$(1.0.1b) \quad x^2 = p^2 + r \sin \theta \sin \phi,$$

$$(1.0.1c) \quad x^3 = p^3 + r \cos \theta.$$

Also recall that the integration measure associated to $B_r(0)$ is $d\sigma = r^2 d\omega$, where $d\omega \stackrel{\text{def}}{=} \sin \theta d\theta d\phi$. Here, ω represents the angular variables. We will abuse notation by using the symbol ω to denote both the angular coordinates (θ, ϕ) , and alternatively as the corresponding point $(\sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta) \in \partial B_1(0)$.

Proposition 1.0.1 (Spherical averages). *Let $u(t, x) \in C^2([0, \infty) \times \mathbb{R}^3)$ be a solution to the 1 + 3 dimensional global Cauchy problem*

$$(1.0.2a) \quad -\partial_t^2 u(t, x) + \Delta u(t, x) = 0, \quad (t, x) \in [0, \infty) \times \mathbb{R}^3,$$

$$(1.0.2b) \quad u(0, x) = f(x), \quad x \in \mathbb{R}^3,$$

$$(1.0.2c) \quad \partial_t u(0, x) = g(x), \quad x \in \mathbb{R}^3.$$

For each $r > 0$, define the spherically averaged quantities

$$(1.0.3a) \quad U(t, r; x) \stackrel{\text{def}}{=} \frac{1}{4\pi r^2} \int_{\partial B_r(x)} u(t, \sigma) d\sigma = \frac{1}{4\pi} \int_{\omega \in \partial B_1(0)} u(t, x + r\omega) d\omega,$$

$$(1.0.3b) \quad F(r; x) \stackrel{\text{def}}{=} \frac{1}{4\pi r^2} \int_{\partial B_r(x)} f(\sigma) d\sigma,$$

$$(1.0.3c) \quad G(r; x) \stackrel{\text{def}}{=} \frac{1}{4\pi r^2} \int_{\partial B_r(x)} g(\sigma) d\sigma,$$

and their related modifications

$$(1.0.4a) \quad \tilde{U}(t, r; x) \stackrel{\text{def}}{=} rU(t, r; x),$$

$$(1.0.4b) \quad \tilde{F}(r; x) \stackrel{\text{def}}{=} rF(r; x),$$

$$(1.0.4c) \quad \tilde{G}(r; x) \stackrel{\text{def}}{=} rG(r; x).$$

Then $\tilde{U}(t, r; x) \in C^2([0, \infty) \times [0, \infty))$ is a solution to the following initial + boundary-value problem for the **one-dimensional** wave equation:

$$(1.0.5a) \quad -\partial_t^2 \tilde{U}(t, r; x) + \partial_r^2 \tilde{U}(t, r; x) = 0, \quad (t, r) \in [0, \infty) \times [0, \infty),$$

$$(1.0.5b) \quad \tilde{U}(t, 0; x) = 0, \quad t \in [0, \infty),$$

$$(1.0.5c) \quad \tilde{U}(0, r; x) = \tilde{F}(r; x), \quad r \in (0, \infty),$$

$$(1.0.5d) \quad \partial_t \tilde{U}(0, r; x) = \tilde{G}(r; x), \quad r \in (0, \infty).$$

Furthermore,

$$(1.0.6) \quad \lim_{r \rightarrow 0} U(t, r; x) = u(t, x).$$

Proof. Differentiating under the integral on the right-hand side of (1.0.3a), using the chain rule relation $\partial_r [u(t, x + r\omega)] d\omega = (\nabla u)(t, x + r\omega) \cdot \omega d\omega = \frac{1}{r^2} \nabla_{\hat{N}(\sigma)} u(t, \sigma) d\sigma$ (where $\hat{N}(\sigma)$ is the outward unit normal to $\partial B_r(x)$), and applying the divergence theorem, we compute that

$$(1.0.7) \quad \partial_r U = \frac{1}{4\pi r^2} \int_{\partial B_r(x)} \nabla_{\hat{N}(\sigma)} u(t, \sigma) d\sigma = \frac{1}{4\pi r^2} \int_{B_r(x)} \Delta_y u(t, y) d^3 y.$$

We now derive a version of the fundamental theorem of calculus that will be used in our analysis below. If h is a continuous function on \mathbb{R}^3 , then using spherical coordinates (ρ, ω) centered at the fixed point x , we have

$$(1.0.8) \quad \partial_r \int_{B_r(x)} h(y) d^3 y = \partial_r \int_0^r \int_{\omega \in \partial B_1(0)} \rho^2 h(\rho, x + \rho\omega) d\omega d\rho = \int_{\omega \in \partial B_1(0)} r^2 h(r, x + r\omega) d\omega \stackrel{\text{def}}{=} \int_{\partial B_r(x)} h(\sigma) d\sigma.$$

Multiplying both sides of (1.0.7) by r^2 and applying (1.0.8), we have that

$$(1.0.9) \quad \partial_r(r^2 \partial_r U) = \frac{1}{4\pi} \partial_r \int_{B_r(x)} \Delta_y u(t, y) d^3 y = \frac{1}{4\pi} \int_{\partial B_r(x)} \Delta u(t, \sigma) d\sigma.$$

Differentiating under the integral in (1.0.3a) and using (1.0.2a), we have that

$$(1.0.10) \quad \partial_t^2 U(t, r; x) = \frac{1}{4\pi r^2} \int_{\partial B_r(x)} \partial_t^2 u(t, \sigma) d\sigma = \frac{1}{4\pi r^2} \int_{\partial B_r(x)} \Delta u(t, \sigma) d\sigma.$$

Comparing (1.0.9) and (1.0.10), we see that

$$(1.0.11) \quad \partial_t^2 U(t, r; x) = \frac{1}{r^2} \partial_r(r^2 \partial_r U) = \partial_r^2 U(t, r; x) + \frac{2}{r} \partial_r U(t, r; x).$$

Multiplying both sides of (1.0.11) by r and performing simple calculations, we see that

$$(1.0.12) \quad \partial_t^2 [rU(t, r; x)] = \partial_r^2 [rU(t, r; x)].$$

We have thus shown that the PDE (1.0.5a) is verified by $\tilde{U} \stackrel{\text{def}}{=} rU$.

Using (1.0.2b) - (1.0.2c) and definitions (1.0.3b) - (1.0.3c), it is easy to check that the initial conditions (1.0.5c) - (1.0.5d) hold. Note that you will have to differentiate under the integral in (1.0.3a) in order to show that (1.0.5d) holds.

The limit (1.0.6) follows easily from the right-hand side of (1.0.3a), since u is continuous.

Finally, the boundary condition (1.0.5b) then follows easily from multiplying (1.0.6) by r before taking the limit $r \rightarrow 0^+$.

□

Corollary 1.0.2 (Representation formula for $\tilde{U}(t, r; x)$). *Under the assumptions of Proposition 1.0.1, for $0 \leq r \leq t$, we have that*

$$(1.0.13) \quad \tilde{U}(t, r; x) \stackrel{\text{def}}{=} rU(t, r; x) = \frac{1}{2} \left(\tilde{F}(r+t; x) - \tilde{F}(r-t; x) \right) + \frac{1}{2} \int_{\rho=-r+t}^{\rho=r+t} \tilde{G}(\rho; x) d\rho.$$

Proof. (1.0.13) follows from (1.0.5a) - (1.0.5d) and the Corollary to d'Alembert's formula.

□

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